information about binding energies and to make isotopic assignments of resonances. The sum-coincidence spectra yield the results  $6.16 \pm 0.05$  and  $5.44 \pm 0.09$  MeV for the binding energies of W<sup>182</sup> (21.2 eV) and W<sup>186</sup> (18.8 eV), respectively; these values are not in disagreement with the values  $6.29 \pm 0.04$  and  $\approx 4.8$  MeV obtained in other ways.40 The maximum energy of the sum pulses observed in the spectrum for the 15.9-eV resonance is  $6.30 \pm 0.08$  MeV. This energy implies that the resonance must be assigned to either W<sup>180</sup> or W<sup>182</sup>. Similarly, a previously reported weak resonance at 75 eV must be assigned to either W<sup>180</sup> or W<sup>182</sup>.

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# Atomic Masses from Gallium to Molybdenum\*

RICHARD R. RIES,<sup>†</sup> RICHARD A. DAMEROW,<sup>‡</sup> AND WALTER H. JOHNSON, JR. School of Physics, University of Minnesota, Minneapolis, Minnesota (Received 14 June 1963)

The 16-in. double-focusing mass spectrometer at the University of Minnesota has been employed to measure the atomic mass of 42 stable isotopes in the region A = 69 to 100. The standard error associated with these results is approximately 5 parts in 108. Improvements in the instrument are described that result in an increase of useful resolution of a factor of 2 to 3. A set of 64 radioactive masses is calculated from the stable mass data together with  $\beta$ -decay energies and nuclear reaction Q values. The resultant table of masses is used to calculate total nuclear binding energies, separation energies and pairing energies for a number of nuclei in the region near N = 50. The systematics of the separation energies display very smooth characteristics except at the shell closure. Neutron pairing energies show a marked decrease in value following the shell closure.

## INTRODUCTION

**HE** 16-in. double-focusing mass spectrometer at the University of Minnesota has been employed in the past to measure atomic masses in several regions of the periodic table. Measurements of atomic masses for most of the stable isotopes with A < 70 have been reported.<sup>1-5</sup> In addition, krypton and xenon,<sup>6</sup> lead and mercury<sup>7</sup> masses have been measured. Operational difficulties in the mass spectrometer become progressively more apparent in measurements of heavy isotopes where maximum resolution is required. These difficulties necessitated the movement and reconstruction of the instrument.

The improved instrument has been employed to measure a number of stable masses for A > 70. Mass doublets in the region from gallium through molyb-

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<sup>4</sup> K. S. Quisenberry, T. T. Scolman, and A. O. Nier, Phys. Rev. 104, 461 (1956).
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<sup>6</sup> H. F. Durbroorth in Prove 11 (1977).

denum are reported in this paper and the following paper<sup>8</sup> reports results in the region from ruthenium through xenon.

The mass results have been employed to study the nuclear binding energy systematics in the region Z=31to 42 and N=36 to 58. These data include the N=50shell closure and also the possible N = 40 and Z = 40 subshell closure.

#### THE INSTRUMENT

The instrument employed for all previous measurements has been described in some detail elsewhere.<sup>9</sup> The ion optics of the instrument are arranged to yield firstand second-order angle focusing and first-order velocity focusing at the fixed collector slit. Doublets are measured by the peak-matching method.

One of the recurring difficulties in the original instrument was the random modulation of the ion beam resulting from building vibrations and time-varying magnetic fields from nearby ac power lines. The modulation not only limited the maximum useful resolution of the mass spectrometer, but also limited the sweep frequency that could be employed.

To remove these difficulties, the instrument was rebuilt in a ground-floor room which had lower stray magnetic fields. The instrument was mounted on a 2-ton cast-iron surface plate. The use of the surface plate as an

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<sup>†</sup> Present address: Max Planck Institut für Chemie, Mainz, Germany. <sup>‡</sup> Present address: Sandia Corporation, Albuquerque, New

<sup>&</sup>lt;sup>1</sup> K. S. Quisenberry, T. T. Scolman, and A. O. Nier, Phys. Rev. 102, 1071 (1956).
<sup>2</sup> T. T. Scolman, K. S. Quisenberry, and A. O. Nier, Phys. Rev. 102 (1975) (1976). 102, 1076 (1956).

<sup>&</sup>lt;sup>6</sup> H. E. Duckworth, in Proceedings of the International Conference on Nuclidic Masses (University of Toronto Press, Hamilton, 1960),

Chap. 7, p. 446. 7 J. L. Benson, R. A. Damerow, and R. R. Ries, Phys. Rev. 113, 1105 (1959),

<sup>&</sup>lt;sup>8</sup> R. A. Damerow, R. R. Ries, and W. H. Johnson, following paper, Phys. Rev. **132**, 1673 (1963). <sup>9</sup> H. Hintenberger, Nuclear Masses and Their Determination

<sup>(</sup>Pergamon Press, Inc., London, 1957), Session VII, p. 185.

accurately plane reference surface greatly aided in the precise location of the various elements of the spectrometer. The spectrometer and surface plate are supported on four spring mounts, the resonant frequency of the supported mass being about 2 cps. This system then effectively isolates the spectrometer from the higher building vibration frequencies. In order to further minimize the effects of vibration, the spectrometer housing is mounted rigidly to the surface plate. Focusing adjustments are now made by moving the magnet on a track system. The modification also included redesign of several of the adjustable slits so as to give more precise control of the ion beam.

The changes made in the instrument have resulted in improved performance. The maximum usable resolution of the improved instrument is 2 or 3 times that of the original spectrometer. Other changes have considerably decreased the time necessary for focusing the instrument.

#### MEASUREMENTS

The mass spectrometer has the property that the mass of the ion collected is proportional to the resistance of a circuit element which determines the electric field in the instrument. Thus, one can show that the equation which relates doublet width to resistance is

$$\Delta M = M \Delta R / R \,, \tag{1}$$

where  $\Delta M$  is the doublet width, M is the mass of one member of the doublet which corresponds to resistance value R, and  $\Delta R$  is the change in resistance necessary to cause the second ion group to be collected. By measuring known doublets, Eq. (1) may be shown to be correct to a high degree of accuracy.

In the past, this instrument was employed mainly to measure narrow doublets made up of a hydrocarbon comparison ion of known mass and an ion of unknown mass at the same mass number. For the mass region under consideration in this paper, these doublets have  $\Delta M/M$  values of approximately 10<sup>-3</sup>. It has been shown that the instrument is capable of measuring wider doublets with sufficient accuracy to be worthwhile.6 One type of wide doublet is the doublet composed of two adjacent isotopes of the same element. Doublets of this sort are known as isotopic doublets and have been employed previously to supplement the hydrocarbon doublet data.

One of the difficulties that arises when hydrocarbon comparison ions are employed is caused by the rare isotope of carbon,  $C^{13}$ . When the hydrocarbon  $C_m^{12}H_n^1$ is used as a comparison ion, a satellite ion is also present,  $C^{13}C_{m-1}{}^{12}H_{n-1}{}^{1}$ . In the mass region considered in this paper, a resolution of about  $1/50\ 000$  is required to completely resolve the C13 satellite. Resolution of this instrument is defined as  $\Delta M/M$  where  $\Delta M$  corresponds to the width at half-height of an ion peak of mass M. The resolution values ranged from  $1/60\,000$  to  $1/200\,000$  during this period. Thus, the C<sup>13</sup> satellite was at all times completely resolved.

TABLE I. Mass doublets.

	Mass	
	difference <sup>b</sup>	
Doublet <sup>a</sup>	(mu)	Error
C5H9-Ga69	144.852 7	24
$C_{5}H_{11} - Ga^{71}$	161.370 2	32
$C_{5}H_{10} - Ge^{70}$	154.001 3	22
$C_4H_8O - Ge^{70}H_9$	117.616 1	18
$C_4H_8O-Ge^{72}$	135.438 4	21
$C_4H_9O-Ge^{73}$	141.878 4	21
$CS_2 - Ge^{74}H_2$	7.314 0	14
$CS_2 - Ge^{76}$ $Ge^{70}H_2 - Ge^{72}$	22.741 6	15
${ m Ge^{70}H_2 - Ge^{72}}$	17.821 3	17
Ge <sup>72</sup> H – Ge <sup>73</sup>	6.443 9	13
${ m Ge^{73}H_2} - { m Ge^{74}H}$	10.105 1	17
$Ge^{74}H_2 - Ge^{76}$	15.425 0	17
$\begin{array}{c} C_{3}H_{7}O_{2}-As^{75}\\ C_{6}H_{2}-Se^{74}\\ C_{6}H_{4}-Se^{76}\\ \end{array}$	123.009 8	26
$C_6H_2 - Se^{r_4}$	93.173 8	38
$C_{6}H_{4} - Se^{76}$	112.099 9	81
$C_6H_5 - Se^{77}$ $C_6H_6 - Se^{78}$	119.211 9 129.642 6	$42 \\ 22$
$C_6H_6 - Se^{80}$ $C_6H_8 - Se^{80}$	129.042 0	22
$C H S_{6}^{82}$	161.545 0	29 46
$C_{6}H_{10} = 56^{-1}$	136.444 3	24
$C_6H_8 - HBr^{81}$	154.135 3	38
$\begin{array}{c} C_{6} B_{10} - S e^{82} \\ C_{6} H_{10} - S e^{82} \\ C_{6} H_{8} - H B r^{79} \\ C_{6} H_{10} - H B r^{81} \\ B r^{81} - H B r^{79} \end{array}$	990.135 1	125
$C_{c}H_{c}-Kr^{78}$	126.584 3 <sup>d</sup>	36
${f C_6 H_6 - Kr^{78} \over C_6 H_8 - Kr^{80}}$	146.225 7 <sup>d</sup>	46
$C H = Kr^{82}$	164.769 8 <sup>d</sup>	34
$C_6H_{11} - Kr^{83}$	171.946 8 <sup>d</sup>	34
$C_{6}H_{12} - Kr^{84}$	182.399 4 <sup>d</sup>	25
$C_6H_{14} - Kr^{86}$	198.936 7 <sup>d</sup>	27
$C_{6}H_{11} - Kr^{83}$ $C_{6}H_{12} - Kr^{84}$ $C_{6}H_{12} - Kr^{84}$ $C_{6}H_{14} - Kr^{86}$ $Kr^{83} - Kr^{82}$	1000.647 9 <sup>d</sup>	120
$Kr^{84} - Kr^{83}$	997.371 6 <sup>d</sup>	120
$C_{6}H_{13} - Rb^{85}$	189.927 6	39
$C_4H_7O_2 - Rb^{87} C_6H_{12} - Sr^{84}$	135.417 8	27
$C_6H_{12} - Sr^{64}$ $C_6H_{14} - Sr^{86}$	$\frac{180.470\ 8}{200.264\ 9}$	26 36
CHO Sr87	135.722 2	35
$\begin{array}{c} C_{4}H_{7}O_{2}-Sr^{87}\\ C_{4}H_{8}O_{2}-Sr^{88}\\ Sr^{87}-Sr^{86} \end{array}$	146.789 1	33 47
$Sr^{87} - Sr^{86}$	999.618 1	115
$Sr^{88} - Sr^{87}$	996.739 6	116
$C_7H_5 - V^{89}$	133.247 0	34
$C_4H_{10}O_2 - Zr^{90}$	163.377 1	55
$C_7 H_7 - Zr^{91}$	149.143 1	44
$C_{r}H_{r}-Zr^{92}$	157.5694	38
$\begin{array}{c} C_{7}H_{10} - Zr^{94} \\ C_{7}H_{12} - Zr^{94} \\ C_{7}H_{12} - Zr^{96} \\ Zr^{91} - Zr^{90} \\ Zr^{92} - Zr^{91} \end{array}$	171.929 4	39
$C_7H_{12} - Zr^{96}$	185.628 0	57
$Zr^{91} - Zr^{90}$	1000.942 0	116
$Zr^{92} - Zr^{91}$	999.397 2	117
$C_7H_9 - Nb^{93}$ $C_7H_8 - Mo^{92}$	164.046 9	35
$C_7H_8 - MO^{22}$	155.790 0	32 32
$C_7H_{10} - MO^{5x}$	173.159 6 180.236 5	32
$C_7H_{10} - Mo^{94}$ $C_7H_{11} - Mo^{95}$ $C_7H_{12} - Mo^{96}$	180.230 5	30
$C_5H_5O_2 - Mo^{97}$	122.937 6	23
$C_5H_5O_2 - Mo^{98}$ $C_5H_6O_2 - Mo^{98}$	131.375.4	23
$C_7H_{16} - Mo^{100}$	217.730 3	42
$C_7H_{16} - Mo^{100} Mo^{95} - Mo^{94}$	1000.757 2	122
$Mo^{96} - Mo^{95}$	998.838 5	124
$Mo^{97} - Mo^{96}$	1001.346 3	123
Mo <sup>98</sup> -Mo <sup>97</sup>	999.386 0	121

<sup>&</sup>lt;sup>a</sup> Throughout this paper C, H, O, and S refer to C<sup>12</sup>, H<sup>1</sup>, O<sup>16</sup>, and S<sup>32</sup>,

<sup>&</sup>lt;sup>a</sup> Infougnout this paper C, H, O, and S tele to C, H, O, and S televise b Mass differences are given in milliunits. All masses and mass differences in this work are measured in a scale in which the atomic mass of C<sup>12</sup> is exactly equal to 12 units (symbol u). The symbol mu refers to one milliunit on the unified mass scale. The symbol  $\mu$ u refers to one microunit. <sup>e</sup> Throughout this paper the errors refer to the last significant figure of the particular result. The errors given in this table are taken from the original experimental data. The resulting error in an atomic mass calculation will be conded off to the parent ui.

will be rounded off to the nearest  $\mu$ u. <sup>d</sup> For completeness, the krypton doublets have been included. They are taken from the M.S. thesis of R. R. Ries, University of Minnesota, 1959 (unpublished).

1664

Metal ions for these measurements were obtained from a variety of sources. In two cases, gases were employed. These were hydrogen bromide and germanium tetrahydride. Pure metals were used for gallium, arsenic, and strontium. In the remaining cases, the following metallic compounds were employed: selenium dioxide, rubidium chloride, yttrium chloride, zirconium tetrachloride, niobium pentachloride, and molybdenum trioxide. Metals and metallic compounds were heated when necessary to obtain sufficient vapor pressure. Sample heating took place in a source furnace essentially the same as that previously described.<sup>3</sup> Two baffles have been added between the sample and the ionization region to improve the vapor distribution.

### RESULTS

The experimental mass doublet differences are presented in Table I. Both the narrow hydrocarbon-isotope doublets and the wide isotopic doublets are included in this table. The mass scale employed is the unified mass scale adopted by the International Union of Pure and Applied Physics in 1960. In this scale, the atomic mass of C<sup>12</sup> is exactly 12 units (symbol u). The errors listed are standard errors and refer to the last quoted figure. These errors are calculated from resistor uncertainties and the standard error of the mean of the original data. The details of the resistor error assignments may be found in an article by Quisenberry et al.4

The doublet values of Table I are combined with the atomic masses of certain secondary mass standards in order to calculate atomic masses. Table II lists the atomic masses of the stable atoms in the mass region  $69 \leq A \leq 100$ . The secondary mass standards which were employed to calculate atomic masses from the doublet data are recorded in Table III. In some cases, the mass of a particular atom was overdeterminded by measurement of isotopic doublets in addition to the usual narrow hydrocarbon-isotope doublets. In these cases, a weighted least-squares adjustment was carried out on the data, and the best values found in this process are recorded as the results in Table II. The error associated with an atomic mass is the square root of the sum of the square of the error on the appropriate doublet value and the square of the error in the hydrocarbon reference mass. Whenever the least-squares adjustment was possible, the error on these atomic masses is just the error resulting from the least-squares calculation.

The final column of Table II provides a direct comparison between the present and other previous mass spectroscopic values. For comparison purposes, mass values were calculated from previous doublet results using the standard masses listed in Table III. The first half of this column contains the former Minnesota values determined by Collins et al.10 on a smaller, less

TABLE II. Atomic masses computed from the doublet data of Table I together with a comparison with previous mass spectroscopic values.

	Present	resultª	1961 Mass	table <sup>b</sup>	Other resu	lts°
Isotope	u	Error	u	Error	u	Erroi
Ga <sup>69</sup>	68.925 5		68.925 682	28	68.925 72ª	4
Ga71	70.924 7		70.924 840	50	$70.924$ $82^{d}$	8
Ge <sup>70</sup>	69.924 2		69.924 277	20	$69.924\ 00^{d}$	6
Ge <sup>72</sup>	71.922 0		71.921 740	50	$71.921 60^{d}$	5 3
Ge <sup>73</sup>	72.923 4		72.923 360	70	$72.923 \ 34^{d}$	3
Ge <sup>74</sup>	73.921 1	77 2	73.921 150	60	$73.921 \ 00^{d}$	6
$Ge^{76}$	75.921 4	02 2	75.921 360	90	75.921 28 <sup>d</sup>	4
$As^{75}$	74.921 5		74.921 580	50	$74.921$ $72^{d}$	4
Se <sup>74</sup>	73.922 4		73.922 450	60	$73.922 54^{d}$	7
Se <sup>76</sup>	75.919 1		75.919 229	48	75.919 27 <sup>d</sup>	4
Se <sup>77</sup>	76.919 9		76.919 934	48		-
Se <sup>78</sup>	77.917 3		77.917 348	$\frac{10}{48}$		
Se <sup>80</sup>	79.916 5		79.916 512	17	79.916 47ª	4
Se <sup>82</sup>	81.916 7	$\frac{1}{102}$ 5	81.916 660	70	81.916 64 <sup>d</sup>	4
Br <sup>79</sup>	78.918 3		78.918 348	19	78.918 40 <sup>d</sup>	$\overline{5}$
Br <sup>81</sup>	80.916 2	28 3	80.916 344	37	80.916 42 <sup>d</sup>	5
Kr <sup>78</sup>	77.920 3		77.920 368	5	77.920 19 <sup>d</sup>	8
Kr <sup>80</sup>	79.916 3		79.916 388	13	77.920 194	0
Kr <sup>82</sup>	81.913 4				81.913 46 <sup>d</sup>	6
			81.913 483	8		6
Kr <sup>83</sup>	82.914 1		82.914 131	8	82.914 06 <sup>d</sup>	5
Kr <sup>84</sup>	83.911 4		83.911 504	5	83.911 51 <sup>d</sup>	5
Kr <sup>86</sup>	85.910 6		85.910 617	8	85.910 80 <sup>d</sup>	6
Rb <sup>85</sup>	84.911 7		84.911 710	60	84.912 03 <sup>d</sup>	6
Rb87	86.909 1		86.909 180	80	86.909 31 <sup>d</sup>	17
Sr <sup>84</sup>	83.913 4	25 3	83.913 376	11	83.913 399 <sup>f</sup>	17
					83.913 25 <sup>d</sup>	15
Sr <sup>86</sup>	85.909 2	78 4	85.909 260	80	85.909 156 <sup>f</sup>	23
					85.909 36 <sup>d</sup>	10
					85.909 357s	
$Sr^{87}$	86.908 8	82 4	86.908 890	80	86.908 816 <sup>f</sup>	42
					86.908 99 <sup>d</sup>	6
Sr <sup>88</sup>	87.905 6	34 5	87.905 610	90	87.905 485	16
					87.905 678s	18
					87.906 01 <sup>d</sup>	11
$\mathrm{Y}^{89}$	88.905 8	76 4	88.905 430	90	88,905 862f	15
~					88,905 72 <sup>d</sup>	11
Zr <sup>90</sup>	89.904 6	96 5	89.904 320	90	89.904 672 <sup>f</sup>	
21	07.7010	000	07.701 020	20	89.904 931	
					89.904 33 <sup>d</sup>	25
$Zr^{91}$	90.905 6	31 4	90.905 250	100	90,905 629 <sup>f</sup>	20
$Zr^{92}$	90.905 0		90.903 230	110	91.905 023 <sup>f</sup>	20
$Zr^{94}$	93.905 0		93.906 140	360	93.906 268 <sup>f</sup>	25
$Zr^{96}$			95.908 200	800	95.900 208 95.908 379 <sup>f</sup>	46
	95.908 2				92.906 315 <sup>f</sup>	25
Nb <sup>93</sup>	92.906 3	75 4	92.906 020	110		
3 6 00	04 00 6 0	0 <b>5</b> 4	01.004.000	420	92.905 66 <sup>d</sup>	8
Mo <sup>92</sup>	91.906 8		91.906 290	130	91.906 869 <sup>f</sup>	42
Mo <sup>94</sup>	93.905 0		93.904 740	130	93.905 166 <sup>t</sup>	52
$Mo^{95}$	94.905 8		94.905 720	360	94.905 841 <sup>f</sup>	15
$Mo^{96}$	95.904 6	70 3	95.904 550	360	95.904 685 <sup>f</sup>	44
$Mo^{97}$	96.906 0	14 3	96.905 750	400	96.905 952 <sup>f</sup>	30
$Mo^{98}$	97.9054		97.905 510	410	97.905 425 <sup>f</sup>	16
$Mo^{100}$	99.907 4	64 5	99.907 570	490	99.907 543f	35

\* The atomic masses are computed from the doublet values of Table I and the values of the secondary mass standards listed in Table III. <sup>b</sup> See Ref. 12.

<sup>o</sup> In the original references, these results were presented in the old O<sup>16</sup> scale. They have been changed to comply to the new C<sup>12</sup> mass scale. <sup>d</sup> See Ref. 10.

• This result is the weighted average of the two narrow Ge72 doublets listed

<sup>1</sup> Table I. <sup>1</sup> See Ref. 11. <sup>8</sup> N. R. Isenor, R. C. Barber, and H. E. Duckworth, Can. J. Phys. 38, 819 (1960).

precise mass spectrometer. These values have errors approximately 10 times larger than those of the present investigation. No systematic difference seems to be evident for the 28 atomic mass comparisons which can be made, since 14 of the former results are higher and

<sup>&</sup>lt;sup>10</sup> T. L. Collins, W. H. Johnson, and A. O. Nier, Phys. Rev. 94, 398 (1954).

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14 lower than the present work. Because of the comparatively large error of the previous results, the comparison of results is not particularly valuable as a test of the reliability of the present data.

Starting with strontium at A = 84, a comparison is made with the work carried out by Demirkhanov et al.<sup>11</sup> These results have errors that are about 5 times those quoted for the present results. A comparison of the 14 stable isotopic masses from Y<sup>89</sup> to Mo<sup>100</sup> reveals moderate agreement between the present results and the results of Demirkhanov. This order of agreement does not hold in the case of several of the strontium masses (Sr<sup>86</sup> and Sr<sup>88</sup>) where the disagreement between the two results to more than 5 times their combined errors. Smaller disagreements occur at Zr<sup>92</sup>, Zr<sup>96</sup>, and Nb<sup>93</sup>. The reason for these discrepancies is unknown.

The present results are also compared with the values found in the Mass Table due to König et al.<sup>12</sup> The results of König represent "best" values from a least-squares adjustment of much of the available nuclear reaction and disintegration data, together with selected mass spectroscopic information. It should be mentioned that in the mass region from A = 69 to A = 93, the mass results of Collins *et al.*<sup>10</sup> served as the mass spectroscopic input data for the least-squares adjustment of the Mass Table. The comparison of the present results with this table is quite good from A = 73 to A = 88. A particularly encouraging feature of this table is the fact that, with three exceptions, the table always adjusted the older Minnesota results of Collins in the right direction to agree with the more precise, present results. Until recently, precise mass information from A = 88 to A = 100 was not available. This is reflected in the poor agreement of the Mass Table results with the present work in this mass region. The discrepancies are particularly obvious for the zirconium and molybdenum

TABLE III. Secondary mass standards.

	Present va	lue	Refer-
Standard	u	Error	ence
C12	12.000 000 0		a
$H^1$	1.007 824 7	2	b
$n^{ m e}$	$1.008\ 665\ 4$	4	d
$H^2$	2.014 102 2	1	d
$N^{14}$	14.003 073 1	4	е
O <sup>16</sup>	15,994 914 2	5	b
S <sup>32</sup>	31.972 073 8	11	d
Cl35	34.968 853 1	19	f

This is the definition of the C12 mass scale.

• See Ref. 2.

TABLE IV. L	east-squares	adjusted	values for	overdetermined
	able isotopes			

Isotope or isotopic difference	Measured resu u Er	Final adjusted <sup>b</sup> lt <sup>a</sup> result ror u Erro	ence
C . 70	69.924 246	2 69.924 247 2	1 1
Ge <sup>70</sup>			+1
Ge <sup>72</sup>	71.922 073	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	+2
Ge <sup>78</sup>	72.923 458	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$-1 \\ -3$
Ge <sup>74</sup>	73.921 180	3 73.921 177 2	
Ge <sup>76</sup>	75.921 402	3 75.921 402 2	0
$Ge^{72} - Ge^{70}$	1.997 828	2 1.997 828 2	0
$Ge^{73} - Ge^{72}$	1.001 381	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	+1
$Ge^{74} - Ge^{73}$	0.997 720	2 0.997720 2	0
$Ge^{76} - Ge^{74}$	2.000 224	2 2.000 224 3 3 78.918 328 3	0
Br <sup>79</sup>	78.918 328	3 78.918 328 3 4 80.916 287 4	0
Br <sup>81</sup>	80.916 287	4 80.916 287 4 2 1.997 959 5	0
Br <sup>81</sup> -Br <sup>79</sup>			-1
Kr <sup>82</sup>	81.913 477	4 81.913 477 4	0
Kr <sup>83</sup> Kr <sup>84</sup>	82.914 125	4 82.914 125 4	
Kr <sup>83</sup> -Kr <sup>82</sup>	83.911 497	3 83.911 497 3 2 1.000 648 5	0
		2 1.000 648 5	0
$Kr^{84} - Kr^{83}$		2 0.997 372 5	0
Sr <sup>86</sup> Sr <sup>87</sup>	85.909 280	4 85.909 278 4	-2
	86.908 879	4 86.908 882 4 5 87.905 634 5	+3
Sr <sup>88</sup> Sr <sup>87</sup> —Sr <sup>86</sup>	87.905 637	5 87.905 634 5	-3
Sr <sup>87</sup> — Sr <sup>86</sup> Sr <sup>88</sup> — Sr <sup>87</sup>		2 0.999 603 5	-15
		2 0.996 752 6 6 89.904 696 5	+12
Zr <sup>90</sup> Zr <sup>91</sup>	89.904 698		-2
	90.905 630	5 90.905 631 4	+1
$Zr^{92}$	91.905 028	4 91.905 028 4 2 1.000 935 7	$-\frac{0}{7}$
$Zr^{91} - Zr^{90}$			
Zr <sup>92</sup> —Zr <sup>91</sup> Mo <sup>94</sup>			0
	93.905 087	4 93.905 086 4 4 94.905 835 4	$-1_{0}$
Mo <sup>95</sup>	94.905 835		0
Mo <sup>96</sup>	95.904 669		+1
Mo <sup>97</sup>	96.906 014	3 96.906 014 3 3 97.905 401 3	0
Mo <sup>98</sup>	97.905 401		0
$Mo^{95} - Mo^{94}$			8
$Mo^{96} - Mo^{95}$			-4
$Mo^{97} - Mo^{96}$			-1
$Mo^{98} - Mo^{97}$	0.999 386 1	.2 0.999 387 4	+1

<sup>a</sup> Calculated from the doublet data of Table I and the standard masses listed in Table III. These results serve as input data for the least-squares adjustment. <sup>b</sup> These results and errors are the least-squares adjusted values. • Adjusted data minus originally measured results (in microunits).

isotopes, where the mass table values also have very large errors.

Hydrocarbon mass unit doublets of the form  $C_mH_n - C_mH_{n-1}$  were measured concurrently with the doublets of Table I in order to make a continuous check on the validity of the dispersion relation, Eq. (1). The average of these mass unit measurements, taken during the measurement of any one element, was used to calculate a value for the hydrogen atomic mass from Eq. (1); and this result was compared with the accepted value for hydrogen listed in Table III. The ratio of the accepted value of the hydrogen mass to the measured value of the hydrogen mass, as calculated from the mass unit doublet, is called the dispersion constant. Thus, when the mass unit measurements lead to results which agree on the average with the accepted mass of hydrogen, the dispersion constant is just equal to unity, and Eq. (1) is used as it stands. However, any deviation from the accepted value of hydrogen defines a dispersion constant different from unity. All doublets which are measured concurrently with this particular mass unit

<sup>1</sup> See Ref. 3.

<sup>&</sup>lt;sup>11</sup> R. A. Demirkhanov, V. V. Dorokhov, and M. I. Dzkuya, N. M. DERIMKHAROV, V. V. DOTOKDOV, and M. I. Dzkuya, Zh. Eksperim. i Teor. Fiz. 40, 1572 (1961) [translation: Soviet Phys.—JETP 13, 1104 (1961)].
<sup>12</sup> L. A. König, J. H. E. Mattauch, and A. H. Wapstra, Nucl. Phys. 31, 18 (1962).

TABLE V. Comparison of mass spectrometer and nuclear	
reaction mass differences for stable isotopes.	

	D		Nuclear re		Reaction	Refer-
Mass difference	Present r u	Error	result u	Error	employed	ence <sup>b</sup>
Ge73-Ge72	1.001 382	2 2	1.001 687	172	Ge <sup>73</sup> (7,n)Ge <sup>72</sup>	c
Se <sup>77</sup> -Se <sup>76</sup>	1.000 713	39	1.000 703	10	$\mathrm{Se}^{76}(n,\gamma)\mathrm{Se}^{77}$	d
Se78 - Se77	0.997 394	£ 5	0.997 411	15	$\mathrm{Se}^{77}(n,\gamma)\mathrm{Se}^{78}$	d
Kr <sup>84</sup> – Kr <sup>83</sup>	0.997 372	2 5	0.997 317	21	$\mathrm{Kr}^{83}(n,\gamma)\mathrm{Kr}^{84}$	5-2-10, 6
Rb <sup>87</sup> -Sr <sup>87</sup>	0.000 30	15	0.000 293	3	Rb <sup>87</sup> (β <sup>-</sup> )Sr <sup>87</sup>	60-3-61
			0.000 261	32	$Rb^{87}(p,n)Sr^{87}$	е
Sr <sup>87</sup> - Sr <sup>86</sup>	0,999 603	35	0.999 629	19	$\mathrm{Sr}^{86}(n,\gamma)\mathrm{Sr}^{87}$	d
			0.999 540	215	$Sr^{86}(d,p)Sr^{87}$	60-3-62
Sr <sup>88</sup> -Sr <sup>87</sup>	0.996 752	2 6	0.996 737	15	$Sr^{87}(n,\gamma)Sr^{88}$	d
Zr <sup>91</sup> -Zr <sup>90</sup>	1.000 935	5 7	1.000 888	32	$Zr^{90}(d,p)Zr^{91}$	f
Zr92 - Zr91	0.999 393	76	0.999 332	32	$Zr^{91}(d,p)Zr^{92}$	g
			0.999 368	43	$Zr^{91}(n,\gamma)Zr^{92}$	d
$Mo^{96} - Mo^{95}$	0.998 834	4 5	0.998 831	11	$Mo^{95}(n, \gamma) Mo^{95}$	d
$Mo^{97} - Mo^{96}$	1.001 345	54	1.001 436	322	Mo <sup>96</sup> (d,p) Mo <sup>97</sup>	h
			1.001 043	322	$Mo^{97}(\gamma,n)Mo^{96}$	60-6-49
Mo <sup>98</sup> – Mo <sup>97</sup>	0.999 383	74	0.999 772	107	Mo97 (d,p) Mo98	h

\*These results were calculated from the appropriate Q value and the masses listed in Table III.
<sup>b</sup> The series of numbers in this column refer to the page numbers of the Nuclear Data Sheets of Ref. 16.
<sup>c</sup> See Ref. 17.
<sup>d</sup> See Ref. 14.
<sup>e</sup> A. J. Elwyn, H. H. Landon, S. Oleksa, and G. N. Glasoe, Phys. Rev. 112, 1200 (1958).
<sup>f</sup> R. L. Preston, H. J. Martin, Jr., and M. B. Sampson, Phys. Rev. 121, 1741 (1961).
<sup>g</sup> H. J. Martin, Jr., M. B. Sampson, and R. L. Preston, Phys. Rev. 125, 942 (1962). H. J. Martin, Jr., M. B. Sampson, and R. L. Preston, Phys. Rev. 125, (1962). 942 (1962). <sup>h</sup> N. S. Wall, Phys. Rev. 96, 664 (1954).

doublet are then corrected by the appropriate amount. Dispersion constants were determined for each element during this investigation and they ranged from  $0.999967 \pm 2$  to  $1.000015 \pm 3$ . The average value of this constant for the entire block of measurements was  $0.999997\pm 2$ .

The effect of the least-squares adjustment on the data where isotopic masses were overdetermined can be seen in Table IV. The second column lists the results which are calculated directly from the doublet data and which then serve as the input data for the adjustment. The third column contains the results of the least-squares process, while the last column records the difference between the adjusted value and the input data. It can be seen that the atomic masses change at most by 2 or 3 microunits, and these changes are always covered by the original errors. The adjusted errors on the atomic mass values generally remain the same or become one microunit smaller. The isotopic mass differences are generally adjusted by a larger amount, as in the case of strontium where the two isotopic mass differences are changed by 12 and 15 microunits. With the exception of the  $\mathrm{Sr}^{87}\mathrm{-}\mathrm{Sr}^{86}$  doublet, these adjustments are also covered by the original errors.

A comparison of mass differences determined both from mass spectroscopic measurements and from nuclear reaction energies is contained in Table V. This table records mass differences only when nuclear reactions connecting 2 stable atoms are available. The conversion factor between Q values and mass units is taken from Cohen<sup>13</sup> and is given by 1 unit=931.476  $\pm 0.004$  MeV.

The number of comparisons in this mass region which involve only 2 stable atoms is not very large, but the over-all agreement is quite good. The agreement is particularly good for recent precise nuclear reaction results. Particular mention should be made of the excellent agreement between the present results and the  $(n,\gamma)$  reaction results determined by Kinsey and Bartholomew.<sup>14</sup> With the exception of the Sr<sup>87</sup>-Sr<sup>86</sup> mass difference, all of these results agree with the present work within the combined errors. This comparison also verifies the gamma spectrum assignments made by Kinsey and Bartholemew when determining the groundstate Q values.

Nuclear reaction and disintegration energies are combined with the stable mass data of Table II in order to calculate the atomic masses of radioactive atoms. The results of such calculations are recorded in Table VI, which lists a final adopted mass for 65 radioactive atoms in the mass region from A = 67 to A = 100. A schematic view of this mass region, exhibiting the various paths by which the radioactive masses may be calculated from the stable atomic masses, is presented in Fig. 1. Several zinc masses from the 1961 Mass Table<sup>11</sup> were utilized to calculate some radioactive gallium masses.

As is evident from Fig. 1, some radioactive masses may be calculated from the stable masses in only one way, while others may be determined by several reaction paths. When only one such path is available, the final adopted value is just the calculated result, so this entry is included only in the final column. In the 25 cases where the radioactive mass can be calculated in more than one way, the result of each calculation is recorded in the second column. The final adopted value for such a radioactive mass is then the weighted average of the several individual results.

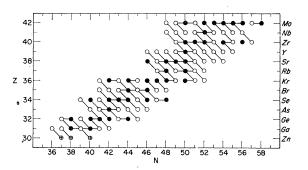


FIG. 1. Reaction scheme showing the stable nuclei whose atomic masses have been measured in this investigation. Solid circles represent stable nuclei, open circles represent radioactive nuclei, and connecting lines indicate available nuclear reaction and and connecting integration and a second and

<sup>&</sup>lt;sup>13</sup> E. R. Cohen, Bull. Am. Phys. Soc. 7, 305 (1962).

<sup>&</sup>lt;sup>14</sup> B. B. Kinsey and G. A. Bartholomew, Can. J. Phys. 31, 1051 (1953).

	Calculated	massª	Nuclear reaction		Final add value			Calculated	massª	Nuclear reaction		Final ado value	
Isotope	u	Error	employed	Referenceb	u	Error	Isotope	e u	Error		Reference <sup>b</sup>	. u	Erro
Ga67			$\operatorname{Zn}^{67}(p,n)\operatorname{Ga}^{67}$	d	66.928 220	12		83.914 561	86	$\mathrm{Rb}^{85}(\gamma,n)\mathrm{Rb}^{84}$	1		
Ga <sup>68</sup>			$Zn^{68}(p,n)Ga^{68}$	d	67.927 996	11	Rb <sup>86</sup>	85.911 186	7	Rb <sup>86</sup> (β <sup>-</sup> )Sr <sup>86</sup>	60-3-50	85.911 186	7
Ga <sup>70</sup>	69.925 936	22	$Ga^{69}(n,\gamma)Ga^{70}$	đ				85.911 173	43	$\mathrm{Rb}^{87}(\gamma,n)\mathrm{Rb}^{86}$	60-3-61		
	69.926 018	11	Ga <sup>70</sup> (β <sup>−</sup> )Ge <sup>70</sup>	đ			Rb <sup>88</sup>	87.911 217	107	Rb <sup>88</sup> (β <sup>-</sup> )Sr <sup>88</sup>	60-3-71	87.911 260	96
	69.926 048	17	Zn <sup>70</sup> (p,n)Ga <sup>70</sup>	d	69.926 012	20		87.911 435	215	Rb <sup>87</sup> (d,p)Rb <sup>88</sup>	60-3-71		
	69.925 956	65	Ga <sup>71</sup> (γ,n)Ga <sup>70</sup>	e			Rb89			Rb <sup>89</sup> (β <sup>−</sup> )Sr <sup>89</sup>	60-3-82	88.911 654	54
Ga <sup>72</sup>	71.926 359	11	Ga <sup>72</sup> (β <sup>-</sup> )Ge <sup>72</sup>	e			Sr <sup>85</sup>	84.912 959	322	$\mathrm{Sr}^{86}(\gamma,n)\mathrm{Sr}^{85}$	60-3-52	84.912 981	33
	71.926 372	9	$Ga^{71}(n,\gamma)Ga^{72}$	ť	71.926 367	7	-	84.912 981	33	Rb <sup>85</sup> ( <i>p</i> , <i>n</i> )Sr <sup>85</sup>	60-3-42		
	71.925 8838	19	$Ga^{71}(n,\gamma)Ga^{72}$	e			Sr <sup>89</sup>	88.907 447	6	Sr <sup>89</sup> (β <sup>-</sup> )Y <sup>89</sup>	60-3-83	88.907 446	9
Ga73			Ga <sup>73</sup> (β <sup>-</sup> )Ge <sup>73</sup>	59-1-37	72.925 123	43		88.907 339	75	Sr <sup>88</sup> ( <i>d</i> , <i>p</i> )Sr <sup>89</sup>	60-3-83		
Ge <sup>69</sup>	68.927 956	5	Ga <sup>69</sup> ( <i>p</i> , <i>n</i> )Ge <sup>69</sup>	h			Sr <sup>90</sup>			Sr <sup>90</sup> (β <sup>-</sup> )Υ <sup>90</sup>	60-4-33	89.907 755	32
	68.927 971	13	Ge <sup>69</sup> (β <sup>+</sup> )Ga <sup>69</sup>	d	68.927 958	5	Sr <sup>91</sup>			Sr <sup>91</sup> (β <sup>-</sup> )Y <sup>91</sup>	60-5-74	90.910 150	16
	68.928 571s		$\mathrm{Ge}^{70}(\gamma,n)\mathrm{Ge}^{69}$	i			Sr <sup>92</sup>			$\mathrm{Sr}^{92}(\beta^-)\mathrm{Y}^{92}$	60-5-84	91.910 949	81
Ge71	70.924 966	32	Ga <sup>71</sup> (p,n)Ge <sup>71</sup>	59-2-44			Y <sup>85</sup>			$Y^{85}(\beta^+)Sr^{85}$	j	84.916 728	37
-	70.924 951	5	Ge <sup>71</sup> (e)Ga <sup>71</sup>	59-2-43	70.924 951	5	Y <sup>86</sup>			$Y^{86}(\beta^+)Sr^{86}$	<b>i</b>	85.914 968	
Ge <sup>75</sup>	74.922 858	22	Ge <sup>75</sup> (β <sup>-</sup> )As <sup>75</sup>	0	74.922 857	22	Y87			$Sr^{87}(p,n)Y^{87}$	60-3-63	86.910 693	215
	74.922 720	215	$Ge^{76}(\gamma,n)Ge^{75}$				Y88	87.909 526	12	$Y^{88}(\beta^+)Sr^{88}$	m		
Ge77			Ge <sup>77</sup> (β <sup>-</sup> )As <sup>77</sup>	59-3-44	76.923 598	55		87.909 329s	33	$Y^{88}(\beta^+)Sr^{88}$	n		
As73			Ge <sup>78</sup> (p,n)As <sup>78</sup>	8	72.923 855	32		87.909 654	86	$\mathrm{Y}^{89}(\gamma,n)\mathrm{Y}^{88}$	1	87.909 522	8
As <sup>74</sup>	73.923 936	11	$As^{74}(\beta^-)Se^{74}$	59-4-71				87.909 901	54	$\mathrm{Y}^{89}(\gamma,n)\mathrm{Y}^{88}$	0		
	73.923 938	31	$\mathrm{As}^{75}(\gamma,n)\mathrm{As}^{74}$	e	73.923 932	8		87.909 519	8	$Sr^{88}(p,n)Y^{88}$	р		
	73.923 928	11	As <sup>74</sup> (β <sup>+</sup> )Ge <sup>74</sup>	e			Y90	89.907 133	12	Y <sup>90</sup> (β <sup>-</sup> )Zr <sup>90</sup>	60-4-34		
$As^{76}$	75.922 387	13	$\mathrm{As^{76}}(\beta^-)\mathrm{Se^{76}}$	59-5-39	75.922 390	12		89.907 189	10	$\mathrm{Y}^{89}(n,\gamma)\mathrm{Y}^{90}$	q	89.907 171	32
	75.922 420	43	$\mathrm{As^{75}}(n,\gamma)\mathrm{As^{76}}$	59-5-41				89.907 420	54	$Y^{89}(d,p)Y^{90}$	r		
As77			$\mathrm{As}^{77}(\beta^-)\mathrm{Se}^{77}$	59-3-46	76.920 646	11	Y <sup>91</sup>			$Y^{91}(\beta^{-})Zr^{91}$	60-5-75	90,907 284	12
As <sup>78</sup>			$\mathrm{As^{78}}(\beta^-)\mathrm{Se^{78}}$	59-5-48, 50	77.921 889	215	Y92			Y <sup>92</sup> (β <sup>-</sup> )Zr <sup>92</sup>	60-5-85	91.908 893	32
A879			As <sup>79</sup> (β <sup>-</sup> )Se <sup>79</sup>	59-2-49	78,920 969	107	$\mathbf{Y}^{93}$			Y <sup>93</sup> (β <sup>-</sup> )Zr <sup>93</sup>	60-5-98	92.909 547	24
Se <sup>78</sup>			Se <sup>73</sup> (β <sup>+</sup> )As <sup>73</sup>	59-1-41	72.926 807	34	Y94			Υ <sup>94</sup> (β <sup>-</sup> )Zr <sup>94</sup>	60-5-11	93.911 685	215
Se <sup>75</sup>			As <sup>75</sup> (p,n)Se <sup>75</sup>	59-1-48	74.922 520	4	Zr <sup>89</sup>	88.908 923	8	Zr <sup>89</sup> (β <sup>+</sup> )Y <sup>89</sup>	60-3-85		
Se <sup>79</sup>			Se <sup>79</sup> (β <sup>-</sup> )Br <sup>79</sup>	59-2-50	78.918 500	6		88.908 678	97	Zr90(γ,n)Zr89	60-4-36	88.908 923	10
Se <sup>81</sup>			$\operatorname{Se}^{\$1}(\beta^{-})\operatorname{Br}^{\$1}$	59-1-64	80.917 790	54		88.908 923	8	Y <sup>89</sup> (p,n)Zr <sup>89</sup>	60-3-85		
Br <sup>76</sup>			Br <sup>76</sup> (β <sup>+</sup> )Se <sup>76</sup>	59-5-43	75.924 127	13	Zr <sup>93</sup>	92.906 443	4	Zr <sup>93</sup> (β <sup>-</sup> )Nb <sup>93</sup>	60-5-99	92.906 444	10
Br77			$Se^{77}(p,n)Br^{77}$	59-3-49	76.921 376	6		92.906 550	43	$Zr^{92}(d,p)Zr^{93}$	60-5-99		
Br <sup>78</sup>	77.921 156	11	$Se^{78}(p,n)Br^{78}$	59-5-52	77.921 156	11	Zr <sup>95</sup>	94.908 097	54	$Zr^{94}(d,p)Zr^{95}$	60-5-127	94.908 037	13
		150	$\operatorname{Br}^{79}(\gamma,n)\operatorname{Br}^{78}$	6				94.908 034	12	Zr <sup>95</sup> (β <sup>-</sup> )Nb <sup>95</sup>	60-5-127		
Br <sup>80</sup>	79.918 519	12	Br <sup>80</sup> (β <sup>-</sup> )Kr <sup>80</sup>	59-1-56			Zr <sup>97</sup>			Zr <sup>97</sup> (β <sup>-</sup> )Nb <sup>97</sup>	60-6-47	96.910 944	33
	79.918 556	7	$\mathrm{Br^{80}}(\beta^+)\mathrm{Se^{80}}$	59-1-56	79.918 544	10	Nb <sup>89</sup>			Nb <sup>89</sup> (\$+)Zr <sup>89</sup>	60-3-86	88.913 088	98
	79.918 496	38	$\operatorname{Br}^{81}(\gamma,n)\operatorname{Br}^{80}$	e			Nb90			Nb <sup>90</sup> (β <sup>+</sup> )Zr <sup>90</sup>	60-4-37	89.911 255	12
	79.918 535	14	$\mathrm{Br}^{79}(n,\gamma)\mathrm{Br}^{80}$	8			Nb <sup>91</sup>			$Mo^{91}(\beta^+)Nb^{91}$	60-5-78	90.906 795	82
Br <sup>82</sup>			Br <sup>82</sup> (β <sup>-</sup> )Kr <sup>82</sup>	59-1-70	81.916 796	4	Nb92	91.906 871	215	$\operatorname{Zr}^{\mathfrak{g2}}(p,n)\operatorname{Nb}^{\mathfrak{g2}}$	e		
Br <sup>83</sup>			Br <sup>83</sup> (β <sup>-</sup> )Kr <sup>83</sup>	59-1-80	82.915 199	22		91.907 186	41	$\mathrm{Nb}^{93}(\gamma,n)\mathrm{Nb}^{92}$	60-5-100	91.907 184	47
Kr77		-	Kr <sup>77</sup> (β <sup>+</sup> )Br <sup>77</sup>	e	76.924 468	22		91.907 325	161	Nb <sup>92</sup> (\$+)Zr <sup>92</sup>	60-5-81		
Kr79	78.920 221		$\operatorname{Kr}^{78}(d,p)\operatorname{Kr}^{79}$	59-2-53	78.920 068	14	Nb <sup>94</sup>	93.907 322	32	$Nb^{93}(n,\gamma)Nb^{94}$	60-5-114		
	78.920 067	6	Kr <sup>79</sup> (β <sup>+</sup> )Br <sup>79</sup>	59-2-52				93.907 253	107	$\mathrm{Nb}^{\mathfrak{g}\mathfrak{g}}(d,p)\mathrm{Nb}^{\mathfrak{g}\mathfrak{g}}$	60-5-114	93.907 314	27
Kr <sup>81</sup>			$\mathrm{Kr}^{80}(d,p)\mathrm{Kr}^{81}$	58-1-39	80.916 605	107		93.907 308	54	Nb <sup>94</sup> (β <sup></sup> )Mo <sup>94</sup>	60-5-113		
Kr <sup>85</sup>	84.912 514	9	Kr <sup>85</sup> (β <sup>-</sup> )Rb <sup>85</sup>	60-3-40	84.912 516	6	$\rm Nb^{95}$			Nb <sup>95</sup> (β <sup></sup> )Mo <sup>95</sup>	60-5-128	94.906 832	6
** **	84.912 519	9	Kr <sup>84</sup> ( <i>d</i> , <i>p</i> )Kr <sup>85</sup>	i CO O CO			$\rm Nb^{97}$			Nb <sup>97</sup> (β <sup>-</sup> )Mo <sup>97</sup>	60-6-48	96.908 088	8
Kr <sup>87</sup>		215	Kr <sup>87</sup> (β <sup>-</sup> )Rb <sup>87</sup>	60-3-60	86.913 363	52	Mo <sup>90</sup>			$\mathrm{Mo}^{90}(\beta^+)\mathrm{Nb}^{90}$	60-4-40	89.913 982	108
	86.913 354	54	$\operatorname{Kr}^{86}(d,p)\operatorname{Kr}^{87}$	j			M0 <sup>91</sup>			$\mathrm{Mo}^{92}(\gamma,n)\mathrm{Mo}^{91}$	60-5-88, 81	90.911 583	75
Kr <sup>88</sup>			Kr <sup>88</sup> (β <sup>-</sup> )Rb <sup>88</sup>	60-3-70	87.914 266	235	Mo <sup>98</sup>	92.906 898	43	$\mathrm{Nb}^{93}(p,n)\mathrm{Mo}^{93}$	60-5-101	92.906 953	70
	83.914 392		Rb <sup>84</sup> (β <sup>-</sup> )Sr <sup>84</sup>	5-2-11				92.907 041	54	$\operatorname{Mo}^{92}(d,p)\operatorname{Mo}^{93}$	60-5-101		
	83.914 363	11	Rb <sup>84</sup> (β <sup>+</sup> )Kr <sup>84</sup>	5-2-11	83.914 374	10	Mo <sup>99</sup>			$\mathrm{Mo^{100}}(\gamma,n)\mathrm{Mo^{99}}$	61-1-66	98.907 495	322
	83.914 411	86	$\mathrm{Rb^{85}}(\gamma,n)\mathrm{Rb^{84}}$	k									

TABLE VI. Atomic masses of radioactive isotopes computed from measured isotopic masses of Table II and available nuclear reaction and beta-decay energies.

For comparison purposes, this column lists the results of the calculations whenever more than one nuclear reaction was available for a particular isotope. No entry is made in this column when only one reaction energy is used. In these cases, the final adopted value is also the calculated value.
 <sup>b</sup> Whenever possible, references are taken from large compilations of nuclear data. References to the original literature are made if that is the only source or if the experimental results lead to calculated masses which differ by large amounts. The sequence of numbers in this column refer to the year, the set, and the page numbers of the Nuclear Data Sheets of Ref. 16.
 <sup>a</sup> The final adopted value is the weighted average of the various calculated results from different reactions. When only one reaction energy is available, the adopted values is the larger of the internal error and the external error as defined in the Nuclear Data Sheets.
 <sup>d</sup> F. Everling, L. A. König, J. H. E. Mattauch, and A. H. Wapstra, Nucl. Phys. 25, 177 (1961).

In 15 of the 25 cases where several reactions were available to determine one radioactive mass, the results agree very well; see, for example, As<sup>74</sup>. When the several L. A. König, J. H. E. Mattauch, and A. H. Wapstra, Nucl. Phys. 28, 1 (1961).
See Ref. 15.
This result is not included in the weighted average.
See Ref. 12.
D. M. Van Patter and Ward Whaling, Rev. Mod. Phys. 29, 757 (1957).
i A. H. Wapstra (private communication).
k See Ref. 17.
See Ref. 18.
m See Ref. 20.
See Ref. 21.
G. A. Bartholomew, P. J. Campion, J. W. Knowles, and G. Manning, Nucl. Phys. 10, 590 (1959).
r N. S. Wall, Phys. Rev. 96, 670 (1954).

independent reaction and disintegration energies combined with different stable atomic masses lead to these similar results, both the energy values and the mass

values are confirmed. In the remaining 10 cases, different reaction paths lead to results that do not agree within the combined error. Several of these cases will now be discussed individually.

The energies of the gamma rays in the complex spectrum resulting from the gallium  $(n,\gamma)$  reaction were measured by Campion et al.,15 but the unambiguous assignment of the ground-state gamma energy of Ga<sup>72</sup> was not possible. A listing of the unassigned gamma rays was given by these authors. An estimate of 7.1 MeV for the last neutron binding energy in Ga<sup>72</sup> was given in the Nuclear Data Sheets.<sup>16</sup> This would suggest that the unassigned 6.971-MeV gamma corresponds to the groundstate reaction. However, this Q value leads to a Ga<sup>72</sup> mass which differs by about 475 keV from the result calculated from the beta decay of Ga<sup>72</sup> to Ge<sup>72</sup>. The level schemes for these 2 isotopes seem to be well known and thus we have chosen the beta disintegration energy in order to calculate the mass of Ga<sup>72</sup>. Assuming that result as correct, we have then assigned the 6.516-MeV gamma as the ground-state gamma for the  $Ga^{71}(n,\gamma)Ga^{72}$ reaction.

The  $(\gamma, n)$  reactions for Rb<sup>84</sup> are available. This reaction is difficult to interpret because of the large nuclear spin change involved. The value by Tobin et al.<sup>17</sup> determined by studying the reaction to the isomeric state of Rb<sup>84</sup> is preferred to the value by Geller *et al.*<sup>18</sup> because of its agreement with the two well-known beta decays of Rb<sup>84</sup>.

The positron decay energy of Y<sup>88</sup> to Sr<sup>88</sup> has been a controversial value for some time. Two distinct groups of results were found for this disintegration energy, and they are best represented by Ramaswamy et al.<sup>19</sup> who quotes the value  $3442\pm31$  keV, and by Rhode et al.<sup>20</sup> who determined the value  $3625 \pm 10$  keV. Two conflicting  $Y^{89}(\gamma, n)Y^{88}$  reaction results were not previously helpful in resolving the discrepancy because of the large error in the mass values for Y<sup>89</sup>. The recent and precise  $Sr^{88}(p,n)Y^{88}$  reaction measured by Shafroth,<sup>21</sup> however, gives a result which agrees closely with the disintegration energy of Rhode. In addition, the present mass determination of Y<sup>89</sup> permits a more precise calculation of  $Y^{88}$  from the two  $(\gamma, n)$  reactions. The mass result calculated from the Q value of Geller *et al.*<sup>18</sup> agrees reasonably well with the results due to Rhode and

Shafroth. The  $(\gamma, n)$  result of Chidley *et al.*<sup>22</sup> is far outside the error of these values, and still further from the result of Ramaswamy's measurement. The final adopted value chosen here for the mass of Y<sup>88</sup> is the weighted average of the three results due to Rhode, Shafroth, and Geller.

In the remaining cases where the agreement between the several possible calculations was not good, evidence did not seem to favor one value over another, so the final adopted value in these cases is just the weighted average.

## NUCLEAR SYSTEMATICS

The atomic masses of stable and radioactive isotopes listed in Table II and Table IV may be employed to calculate the nuclear binding energies for various combinations of nucleons. In order to calculate the total nuclear binding energy (TNBE), the proper atomic mass value is substituted into Eq. (2)

TNBE
$$(Z,N) = ZM_{\rm H} + NM_n - zM_N^A - E_b(Z,N)/c^2$$
, (2)

where  $M_{\rm H}$  is the hydrogen mass,  $M_n$  is the neutron mass,  $_ZM_N{}^A$  is the atomic mass of the atom with Z protons and N neutrons. The term  $E_b(Z,N)$  is an estimate of the total electronic binding energy. The functional form of this relation was taken from Foldy.<sup>23</sup> The value ranged from 51  $\mu$ u for gallium to 104  $\mu$ u for molybdenum, with a stated accuracy of 10%. Consideration of the small binding energy of the electron in hydrogen was neglected in view of the much larger uncertainty in the value of  $E_b(Z,N).$ 

Values of the total nuclear binding energy are useful in the calculation of various binding energy quantities of interest in the study of nuclear systematics. One quantity which indicates the general tendency of nuclear binding is the average binding energy, TNBE(Z,N)/A. The binding energy of a given system of nucleons in a particular nucleus may be found by forming the appropriate difference between two total binding energy terms. As an example, the neutron separation energy and the proton separation energy are found from Eq. (3)and Eq. (4), respectively.

$$S_n(Z,N) = \text{TNBE}(Z,N) - \text{TNBE}(Z,N-1), \quad (3)$$

$$S_p(Z,N) = \text{TNBE}(Z,N) - \text{TNBE}(Z-1,N).$$
 (4)

Table VII lists the total nuclear binding energy, the average nuclear binding energy, the neutron separation energy, and the proton separation energy in nuclei for which these quantities may be calculated from the present mass values.

In Fig. 2, the average nuclear binding energy for stable isotopes is plotted as a function of the mass number. The even-A points for each element are connected by a solid line and the odd-A points for all the elements are connected by a dashed line. The character-

<sup>&</sup>lt;sup>15</sup> P. J. Campion and G. A. Bartholomew, Can. J. Phys. 35, 1361 (1957).

<sup>&</sup>lt;sup>(1977)</sup>. <sup>16</sup> Nuclear Data Sheets, compiled by K. Way et al. (Printing and Publishing Office, National Academy of Sciences–National Re-search Council, Washington 25, D. C., 1958–1961). <sup>17</sup> R. Tobin, J. McElhinney, and L. Cohen, Phys. Rev. 110, 1388

<sup>(1958).</sup> <sup>18</sup> K. N. Geller, J. Halpern, and E. G. Muirhead, Phys. Rev. 118,

<sup>1302 (1960).</sup> <sup>19</sup> M. H. Ramaswamy and P. S. Jastram, Nucl. Phys. 19, 243

<sup>(1960).</sup> <sup>20</sup> J. I. Rhode, O. E. Johnson, and W. G. Smith, Phys. Rev. **129**, 815 (1963). <sup>21</sup> S. M. Shafroth, Nucl. Phys. **28**, 649 (1961).

<sup>&</sup>lt;sup>22</sup> B. G. Chidley, L. Katz, and S. Kowalski, Can J. Phys. 36, 407 (1958). <sup>23</sup> L. L. Foldy, Phys. Rev. 83, 397 (1951).

			TNBE <sup>a</sup>	TNRE	r/4b	Sn	,	S	p <sup>d</sup>				TNBE*	TNRE	//b	Sn	3	S,	
Isotope	Ζ	N	mu			mu n		mu	Error	Isotope	Z	N	mu			mu		mu	Error
Zn <sup>66</sup>	30	36	620.600°							Kr <sup>86</sup>	36	50	804.278	9.3521	3	10.572	10		
$Zn^{67}$	30	37	628.165							Kr <sup>87</sup>	36	51	810.190	9.3125	7	5.912	52		
Zn <sup>68</sup> Zn <sup>69</sup>	30 30	38 39	639.114° 645.992°							Kr <sup>88</sup> Rb <sup>84</sup>	36 37	52 47	817.952 782.337	9.2949 9.3135	27 3	7.762	241	7.571	11
$Zn^{70}$	30	40	655.962°							Rb <sup>85</sup>	37	48	793.583	9.3363		11.246	11	7.523	5
Zn <sup>71</sup>	30	$\tilde{41}$	662.005°							Rb86	37	49	802.856	9.3355	3	9.273	8	9.150	9
$Zn^{72}$	30	42	670.901°							Rb <sup>87</sup>	37	50	813.524	9.3509	3	10.668	8	9.246	
Ga <sup>67</sup>	31	36	626.249	9.3470		0.000		5.649		Rb <sup>88</sup>	37	51	820.112	9.3195		6.588	96	9.922	
Ga <sup>68</sup> Ga <sup>69</sup>	31 31	37 38	$635.138 \\ 646.230$	9.3403 9.3657		8.889 11.092	16 11	6.973 7.116	16 9	Rb <sup>89</sup> Sr <sup>84</sup>	37 38	52 46	$828.384 \\782.440$	9.3077 9.3148	7 3	8.272	110	10.432	241
Ga <sup>70</sup>	31	39	654.453	9.3493		8.223	20	8.461	35	Sr <sup>85</sup>	38	47	791.549	9.3140	5	9.109	33	9.212	34
Ga <sup>71</sup>	31	40	664.430	9.3582		9.977	$\tilde{20}$	8.468	16	Sr <sup>86</sup>	38	48	803.917	9.3479	3	12.368	33	10.334	
Ga <sup>72</sup>	31	41	671.429	9.3254	3	6.999	8	9.424		Sr <sup>87</sup>	38	49	812.979	9.3446	3	9.062	5	10.123	8
Ga <sup>73</sup>	31	42	681.338	9.3334		9.909	44	10.437	224	Sr <sup>88</sup>	38	50	824.892	9.3738		11.913	6	11.368	
Ge <sup>69</sup> Ge <sup>70</sup>	32 32	37 38	642.997	9.3188		10 277	F	7.859 9.144	12	Sr <sup>89</sup>	38	51 52	831.746	9.3455 9.3345	3	6.854	10	11.634	
Ge <sup>71</sup>	32 32	30 39	655.374 663.335	9.3625 9.3427		$12.377 \\ 7.961$	5 5	9.144 8.882	$\frac{3}{21}$	Sr <sup>90</sup> Sr <sup>91</sup>	38 38	52 53	840.102 846.373	9.3343	43	8.356 6.271	33 36	11.718	05
Ge <sup>72</sup>	32	40	674.876	9.3733		11.541	5	10.446	4	Sr <sup>92</sup>	38	54	854.239	9.2852	9	7.866	83		
Ge <sup>73</sup>	32	41	682.160	9.3447	3	7.284	3	10.731	7	Y <sup>85</sup>	39	46	786.957	9.2583	5			4.517	37
Ge <sup>74</sup>	32	42	693.105	9.3663		10.945	3	11.767	43	$\mathbf{Y^{86}}$	39	47	797.382	9.2719		10.425	65	5.833	63
Ge <sup>75</sup>	32	43	700.091	9.3345		6.986	22			Y <sup>87</sup> Y <sup>88</sup>	39	48	810.322	9.3140		12.940	220	6.405	
Ge <sup>76</sup> Ge <sup>77</sup>	32 32	$\frac{44}{45}$	710.212 716.680	9.3449 9.3075		$10.121 \\ 6.468$	22 55			Y 89	39 39	49 50	820.159 832.470	9.3200 9.3536		9.837 12.311	215 9	7.180 7.578	
As <sup>73</sup>	33	$\frac{43}{40}$	680.917	9.307		0.400	55	6.041	32	Y90	39	51	839.841	9.3316		7.371	32	8.095	
As <sup>74</sup>	33	41	689.506	9.3176		8.589	33	7.346		<b>Y</b> <sup>91</sup>	39	52	848.393	9.3230		8.552	34	8.291	34
As <sup>75</sup>	33	42	700.512	9.3402		11.006	9	7.407	3	Y <sup>92</sup>	39	53	855.450	9.2984		7.057	34	9.077	
As <sup>76</sup>	33	43	708.378	9.3208		7.866	12	8.287	25	Y93	39	54	863.461	9.2845		8.011	40	9.222	84
As <sup>77</sup> As <sup>78</sup>	33 33	44 45	718.788 726.210	9.3349 9.3104		$10.410 \\ 7.422$	$\frac{16}{215}$	8.576 9.530		Y <sup>94</sup> Zr <sup>89</sup>	39 40	55 49	869.988 828.578	9.2552 9.3099	23 3	6.527	216	8.419	13
As <sup>79</sup>	33	45 46	735.796	9.3104		9.586	213	9.550	222	Zr <sup>90</sup>	40 40	49 50	841.470	9.3099		12.892	11	9.000	
Se <sup>73</sup>	34	39	677.120	9.2756		2.000	210			Zr <sup>91</sup>	40	51	849.201	9.3319		7.731	7	9,360	
Se <sup>74</sup>	34	40	690.117	9.3259		12.997	34	9.200		Zr <sup>92</sup>	40	52	858.469	9.3312		9.268	6	10.076	
Se <sup>75</sup>	34	41	698.738	9.3165		8.621	6	9.232		Zr <sup>93</sup>	40	53	865.718	9.3088		7.249	11	10.268	
Se <sup>76</sup>	34 34	42 43	710.725	9.3516		11.987	9 9	10.213 10.300		Zr <sup>94</sup> Zr <sup>95</sup>	$\frac{40}{40}$	54 55	874.510 881.456	9.3033 9.2785		8.792 6.946	11 14	11.049	
Se <sup>77</sup> Se <sup>78</sup>	34 34	43 44	718.678 729.949	9.3335 9.3583		7.953 11.271	5	11.161	13 11	Zr <sup>96</sup>	40 40	55 56	889.890	9.2785	3	8.434	14	11.468	215
Se <sup>79</sup>	34	45	737.420	9.3344		7.471	6	11.210		Zr <sup>97</sup>	40	57	895.880	9.2359		5.990	34		
Se <sup>80</sup>	34	46	748.056	9.3507	' 3	10.636	7	12.260		Nb <sup>89</sup>	41	48	823.566	9.2536	11				
$\mathrm{Se}^{81}$	34	47	755.461	9.3267		7.405	54			Nb <sup>90</sup>	41	49	834.064	9.2674		10.498	99	5.486	16
Se <sup>82</sup>	34	48	765.214	9.3319		9.753	54	6 01 2	1.4	Nb <sup>91</sup>	41 41	50 51	847.190 855.466	9.3098 9.2985	9 6	13.126 8.276	83 95	5.720 6.265	
Br <sup>76</sup> Br <sup>77</sup>	35 35	41 42	704.951 716.367	9.2757 9.3035		11.416	14	6.213 5.642		Nb <sup>92</sup> Nb <sup>93</sup>	41 41	51 52	855.400 864.940	9.2985		8.270 9.474	95 47	0.205 6.471	
Br <sup>78</sup>	35	43	725.253	9.2981		8.886	13	6.575		Nb94	41	53	872.667	9.2837	4	7.727	27	6.949	
$\mathrm{Br}^{79}$	35	44	736.746	9.3259	3	11.493	11	6.797	3	Nb <sup>95</sup>	41	54	881.814	9.2823		9.147	28	7.304	. 7
Br <sup>80</sup>	35	45	745.196	9.3150		8.450	10	7.776		$Nb^{96}$	41	55	889.260	9.2631	4	7.446	33	7.804	
$Br^{81}$	35	46	756.118	9.3348		10.922	11	8.062	5	Nb97	41	56	897.889	9.2566		8.629	33	7.999	
Br <sup>82</sup> Br <sup>83</sup>	35 35	47 48	764.274 774.537	9.3204 9.3318		8.156 10.263	6 22	8.813 9.323	54 23	M0 <sup>90</sup> M0 <sup>91</sup>	42 42	48 49	830.491 841.555	9.2277 9.2479		11.064	131	6.925 7.491	146 76
$Kr^{77}$	36	40 41	712.431	9.3310		10.203	44	9.323 7.480		Mo <sup>92</sup>	42	50	854.996	9.2479		13.441	75	7.806	
Kr <sup>78</sup>	36	$\overline{42}$	725.200	9.2974		12.769	22	8.833	<b>7</b>	Mo <sup>93</sup>	42	51	863.516	9.2851	8	8.520	70	8.050	84
Kr <sup>79</sup>	36	43	734.161	9.2932	3	8.961	14	8.908	18	Mo <sup>94</sup>	42	52	874.048	9.2984	3	10.532	70	9.108	
Kr <sup>80</sup>	36	44	746.523	9.3315		12.362	15	9.777	5	Mo <sup>95</sup>	42	53	881.964	9.2838	3	7.916	5	9.297	
Kr <sup>81</sup> Kr <sup>82</sup>	36 36	$\frac{45}{46}$	754.955 766.749	9.3204 9.3506		8.432 11.794	107 107	9.759 10.631	107 5	M0 <sup>96</sup> M0 <sup>97</sup>	42 42	54 55	891.796 899.116	9.2895 9.2692	3 3	9.832 7.320	5 4	9.982 9.856	
Kr <sup>83</sup>	36	40 47	774.766	9.3345		8.017	5	10.031	5	Mo <sup>98</sup>	42	56	908.395	9.2692		9.279	4 4	10.506	
Kr <sup>84</sup>	36	48	786.060	9.3579		11.294	4	11.523	22	Mo <sup>99</sup>	42	57	914.966	9.2421	33	6.571	320	20,000	-
$Kr^{85}$	36	49	793.706			7.646	7			$Mo^{100}$	42	58	923.662	9.2366	3	8.696			

<sup>a</sup> Total nuclear binding energy in milliunits. The mass equivalent of the electronic binding energy has been subtracted from the total atomic binding energy. Errors are not listed because, for most purposes, the differences in 2 TNBE values are employed. An estimate of the error associated with TNBE may be found by multiplying the error assigned to TNBE/A by the corresponding A. <sup>b</sup> Average nuclear binding energy per nucleon in milliunits.

Neutron separation energy in milliunits.
Proton separation energy in milliunits.
The zinc masses were not measured in this investigation. In order to study the nuclear systematics of Ga and Ge it was necessary to calculate TNBE for the listed zinc isotopes. The atomic masses of zinc were taken from the 1961 Mass Table of Ref. 12.

istic parabolic shape in the even-A curves for each element appear in this region as has been observed in other regions. The obvious change in the curve connecting odd-A points near A = 89 is a result of the shell closure at N=50. There is an increase in the odd-A

curve in the region from A = 79 to A = 89. This rise does not appear at the shell closure at  $N=82^{24}$  or  $N=126.^{25}$ 

<sup>24</sup> W. H. Johnson, Jr., and A. O. Nier, Phys. Rev. 105, 1014 (1957). <sup>25</sup> V. B. Bhanot, W. H. Johnson, and A. O. Nier, Phys. Rev. 120, <sup>25</sup> (1997).

235 (1960).

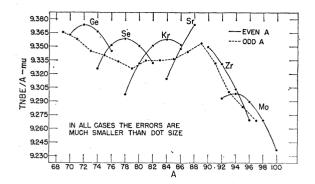


FIG. 2. Average binding energy per nucleon for stable isotopes.

A more detailed study of binding-energy systematics may be made by considering the neutron separation energy and the proton separation energy. Figures 3 and 4 plot the neutron separation energy as a function of the neutron number for even N and odd N, respectively. In each case, successive points for a particular element are connected by a straight line. The sharp discontinuity beyond N=50 is shown in these graphs with greater precision than previously available. The generally smooth character of the curves on both sides of N = 50is perhaps the most significant result inferred from these plots. The N = 42 data points for bromine and the N = 40data point for gallium seem to contradict this smooth behavior. In the case of bromine, the assignment of a mass to the radioactive Br<sup>77</sup> and Br<sup>76</sup> may be in error. In the case of gallium, these is no obvious error in either mass used to calculate this separation energy. This value may indicate an anomaly at N = 40. The  $S_n$  values for germanium and selenium at N=40, however, show no particular anomaly.

Proton separation energies  $S_p$  have been plotted in Fig. 5 for even Z. A line connects data points of constant N. These curves indicate, as has been pointed out previously, that the closure of a neutron shell seems to have no effect on the proton separation energies. The variation of  $S_p$  for a given N value as a function of Z is also smooth. There is a persistent change in slope at Z=40

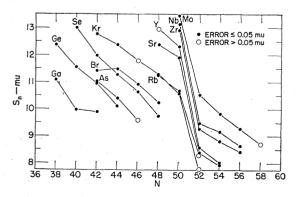


FIG. 3. Neutron separation energies for nuclei with an even number of neutrons,

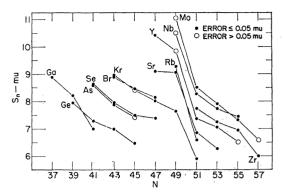


FIG. 4. Neutron separation energies for nuclei with an odd number of neutrons.

for the four curves that have data points at both Z=38 and Z=42. As in the neutron data, this may be an indication of a slight change in nuclear structure near nucleon number 40.

The study of the systematics of the binding energy of the last pair of nucleons in a nucleus is worthwhile because stable masses with small experimental errors are employed in most cases in the calculation. Table VIII lists the value of the binding energy of the last pair of neutrons  $S_{2n}$  for even N nuclei. These data are plotted in Fig. 6 as a function of neutron number N. Values from the same element are connected by a line. Once again the shell closure at N=50 is clearly visible. The smooth behavior of these curves for values on either side of N=50 is clearly evident.

The values of the binding energy of the last pair of protons for even Z nuclei are listed in Table IX. These data are plotted in Fig. 7. Once again the smooth variation of these data with changes in Z is evident. These results also indicate a change in slope at Z=40. Because there are only two other values of N for which there are 3 data points for a given N value, the consistency of slope for values of Z other than 40 is difficult to demonstrate.

A number of pairing energies for neutron and proton pairs may also be calculated from the mass data. The pairing energy for the last pair of neutrons added to a

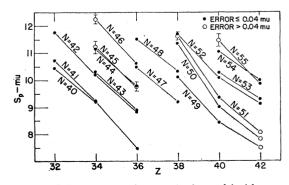


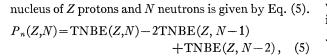
FIG. 5. Proton separation energies for nuclei with an even number of protons.

_	-		$S_{2n}$	a	Ρ,		_	_		$S_{2p}$	,a	$P_{i}$	p <sup>b</sup>
Isotope	Ζ	N	mu	Error	mu	Error	Isotope	Z	Ν	mu	Error	mu	Error
Ga <sup>69</sup>	31	38	19.981	12	2.203	25	Ge <sup>69</sup>	32	37	14.832	12	0.886	25
Ga71	31	40	18.200	4	1.754	40	Ge <sup>70</sup>	32	38	16.260	9	2.028	11
Ga <sup>73</sup>	31	42	16.908	43	2.910	45	Ge71	32	39	17.343	29	0.421	50
Ge <sup>72</sup>	32	40	19.502	3	3.580	10	Ge <sup>72</sup>	32	40	18.914	16	1.978	18
Ge <sup>74</sup>	32	42	18.229	3	3.661	5	Ge <sup>73</sup>	32	41	20.155	220	1.307	220
Ge <sup>76</sup>	32	44	17.107	2	3.135	44	Ge <sup>74</sup>	32	42	22.204	220	1.330	235
$As^{75}$	33	42	19.595	32	2.417	36	Se <sup>73</sup>	34	39	13.785	34		
As <sup>77</sup>	33	44	18.276	11	2.544	27	Se <sup>74</sup>	34	40	15.241	4	3.159	64
As <sup>79</sup>	33	46	17.008	108	2.164	445	Se <sup>75</sup>	34	41	16.578	4	1.886	17
Se <sup>76</sup>	34	42	20.608	9	3.366	12	Se <sup>76</sup>	34	42	17.620	8	2.806	10
Se <sup>78</sup>	34	44	19.224	8	3.318	12	Se <sup>77</sup>	34	43	18.587	22	2.013	33
Se <sup>80</sup>	34	46	18.107	4	3.165	13	Se <sup>78</sup>	34	44	19.737	3	2.585	22
Se <sup>82</sup>	34	48	17.158	6	2.348	108	Se <sup>79</sup>	34	45	20.740	55	1.680	433
Br <sup>79</sup>	35	44	20.379	7	2.607	23	Kr77	36	41	13.693	22	1.267	34
Br <sup>81</sup>	35	46	19.372	5	2.472	21	Kr <sup>78</sup>	36	42	14.475	9	3.191	15
Br <sup>83</sup>	35	48	18.419	22	2.107	24	Kr <sup>79</sup>	36	43	15.483	15	2.333	26
Kr <sup>80</sup>	36	44	21.323	6	3.401	29	Kr <sup>80</sup>	36	44	16.574	5	2.980	7
Kr <sup>82</sup>	36	46	20.226	6	3.362	214	Kr <sup>81</sup>	36	45	17.535	107	1.983	109
Kr <sup>84</sup>	36	48	19.311	4	3.277	8	Kr <sup>82</sup>	36	46	18.693	4	2.569	9
Kr <sup>86</sup>	36	50	18.218	4	2.926	13	Kr <sup>83</sup>	36	47	19.305	54	1.679	55
Kr <sup>88</sup>	36	52	13.674	235	1.850	257	Kr <sup>84</sup>	36	48	20.846	5	2.200	44
$\mathbf{Rb^{87}}$	37	50	19.941	5	1.395	15	Sr <sup>84</sup>	38	46	15.691	4		
Rb <sup>89</sup>	37	52	14.860	54	1.684	200	Sr <sup>85</sup>	38	47	16.783	33	1.641	39
Sr <sup>86</sup>	38	48	21.477	5	3.259	66	Sr <sup>86</sup>	38	48	17.857	4	2.811	9
Sr <sup>88</sup>	38	50	20.975	6	2.851	9	Sr <sup>87</sup>	38	49	19.273	7	0.973	16
Sr <sup>90</sup>	38	52	15.210	32	1.502	37	Sr <sup>88</sup>	38	50	20.614	5	2.122	8
Sr <sup>92</sup>	38	54	14.137	87	1.595	93	Sr <sup>89</sup>	38	51	21.556	53	1.712	199
$Y^{87}$	39	48	23.365	220	2.515	243	Sr <sup>90</sup>	38	52	22.150	237	1.286	259
$\mathbf{Y^{89}}$	39	50	22.148	215	2.474	216	Zr <sup>89</sup>	40	49	15.599	11	1.239	19
$Y^{91}$	39	52	15.923	13	1.181	65	Zr <sup>90</sup>	40	50	16.578	7	1.422	10
$Y^{93}$	39	54	15.068	27	0.954	69	$Zr^{91}$	40	51	17.455	10	1.265	65
Zr <sup>92</sup>	40	52	16.999	7	1.537	11	Zr <sup>92</sup>	40	52	18.367	32	1.785	40
Zr <sup>94</sup>	40	54	16.041	6	1.543	21	Zr <sup>93</sup>	40	53	19.345	19	1.191	67
Zr <sup>96</sup>	40	56	15.380	7	1.488	27	Zr <sup>94</sup>	40	54	20.271	81	1.827	94
$Nb^{91}$	41	50	23.624	128	2.628	130	$Mo^{91}$	42	49	12.977	76	2.005	79
Nb <sup>93</sup>	41	52	17.750	82	1.198	125	$Mo^{92}$	42	50	13.526	6	2.086	164
$\rm Nb^{95}$	41	54	16.874	7	1.420	54	$Mo^{93}$	42	51	14.315	70	1.785	117
$Nb^{97}$	41	56	16.075	10	1.183	65	$Mo^{94}$	42	52	15.579	5	2.637	9
$Mo^{92}$	42	50	24.505	108	2.377	185	$Mo^{95}$	42	53	16.246	11	2.348	55
$Mo^{94}$	42	52	19.052	5	2.012	140	$Mo^{96}$	42	54	17.286	5	2.678	13
Mo <sup>96</sup>	42	54	17.748	4	1.916	8	$Mo^{97}$	42	55	17.660	13	2.052	65
$Mo^{98}$	$\overline{42}$	56	16.599	4	1.959	6	$Mo^{98}$	42	56	18.505	6	2.507	17
Mo <sup>100</sup>	42	58	15.267	5	2.125	644	$Mo^{99}$	42	57	19.086	322		

 
 TABLE VIII. Binding energy and pairing energy of the last pair of neutrons for the listed isotopes.

 
 TABLE IX. Binding energy and pairing energy of the last pair of protons for the listed isotopes.

Binding energy of the last pair of neutrons in milliunits.
 Pairing energy of the last pair of neutrons in milliunits.



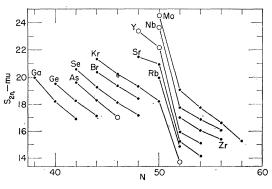


FIG. 6. Binding energy of the last two neutrons.

Binding energy of the last pair of protons in milliunits.
 Pairing energy of the last pair of protons in milliunits.

where N is even. The proton pairing energy  $P_p$  is defined in a similar manner. For light nuclei Mayer and Jensen<sup>26</sup> have concluded that a correlation exists be-

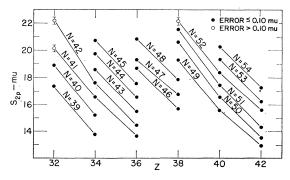


FIG. 7. Binding energies of the last two protons.

<sup>26</sup> M. G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure* (John Wiley & Sons, Inc., New York, 1955).

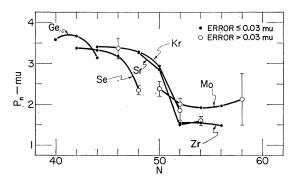


FIG. 8. Neutron pairing energy.

tween the pairing energy and the j value of the odd nucleon of the pair, with larger pairing energies correlated to higher j values. In regions of high j values, one finds that it is energetically possible to have the pair occupy a high j-value state rather than to pair in the lower spin state of the preceding odd nucleon. This mechanism is used to explain the absence of the highest j values from the ground-state spins of odd nuclei.

Neutron and proton pairing energies that may be calculated from the present masses are listed in Tables VIII and IX, respectively. The  $P_n$  values are plotted in Fig. 8 as a function of N. Values from the same element are connected by a curved line. An attempt has been made to correlate the magnitude of these pairing energies with (a) the *j* value of the previous odd neutron and (b) with the j value which the pair is assumed to have according to the filling scheme of Mayer and Jensen.<sup>26</sup> Neither comparison produces very convincing results. The correlation between the value of j and the pairing energy is in some cases what Mayer and Jensen have suggested; in others, the opposite. The one positive statement that may be made is that the value of  $P_n$ decreases, rather strikingly, following the shell closure at N=50. The general tendency of the curves for each element is smooth, and in most instances continuously decreasing with increasing N to N = 50. The character of the curves changes abruptly following N = 50. Beyond

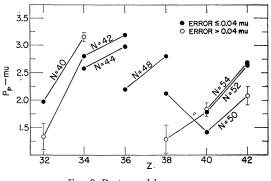


FIG. 9. Proton pairing energy.

N=50, the  $P_n$  values are small, and for each element, are essentially constant for the region plotted. Note that there is nothing anomalous about the one value at N=40. Values of  $P_n$  for zinc from Quisenberry *et al.*<sup>1</sup> at N=40 and N=38 further strengthen this conclusion.

Figure 9 is an illustration of the proton pairing energy  $P_p$  as a function of Z. In this illustration, points with the same N value are connected by lines. An attempt to correlate the  $P_p$  with the j value for the pair is again not particularly fruitful. The  $P_p$  values at Z=40, with  $j=\frac{1}{2}$  for this pair, appear to be lower than practically all other values in this region. A j value of  $\frac{1}{2}$  occurs only at Z=40; values for other pairs in this region are all larger. Thus, in this instance, low  $P_p$  is correlated with low j. There are other instances, however, where this correlation is reversed. It is of interest to note that the value of  $P_p$  at Z=40 is so small. This may indicate a structure change near Z=40 that is not indicated at N=40.

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